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Physicists**



**L'Association canadienne des
physiciens et physiciennes**

The 2014 CAP Undergraduate Lecture Tour

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- Current developments in physics
- Physics research and education in Canada
- the CAP !

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- A strong and effective advocacy group for support of, and excellence in, physics research and education

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Association canadienne des physiciens et physiciennes



Laurentian University
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The 2014 CAP Undergraduate Lecture Tour

Special thanks to our National and Regional Coordinators, the CAP Office, volunteer fund raisers, our sponsors, CAP member/host physics departments, our outstanding CAP volunteer speakers, Canadian physics student societies and
YOU – our audience !

**We hope you enjoy this
event!**

What's the matter with gravity?

Sanjeev Seahra (University of New Brunswick, Fredericton)

CAP Undergraduate Lecture Series: Winter 2014

Successes and failures

Small scale patches

Large scale fixes

Where do we stand?



Gravitation is the most familiar force in our everyday lives, but it remains an active area of physics research. Why is that?

Successes and failures

History of gravity

Testing general
relativity

Shortcomings of GR

How not to quantize
gravity

Is this a problem?

“Observable” quantum
gravity?

Hard to see

Inflation as a Planck
scale microscope

Only testable QG
prediction?

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Quick history of gravitation

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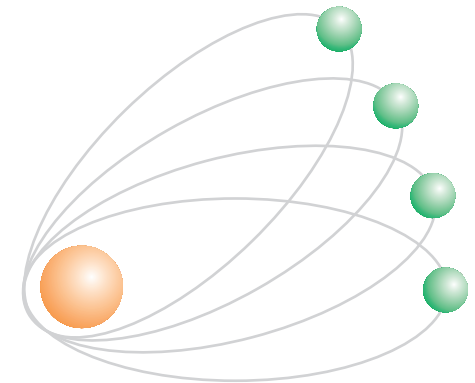
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- **1677:** Newton proposed inverse-square law of gravitation from which he derived Kepler’s laws
- **1800s:** observations of orbit of Mercury were found to be inconsistent with Kepler’s laws
 - the “dark planet” Vulcan was proposed to explain the discrepancy



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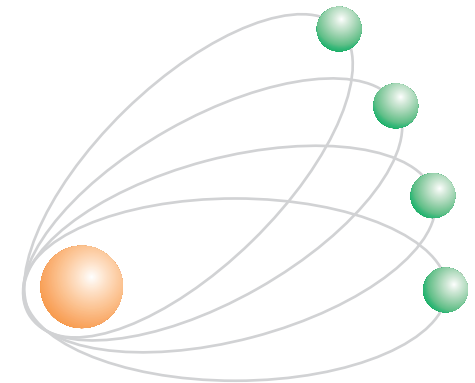
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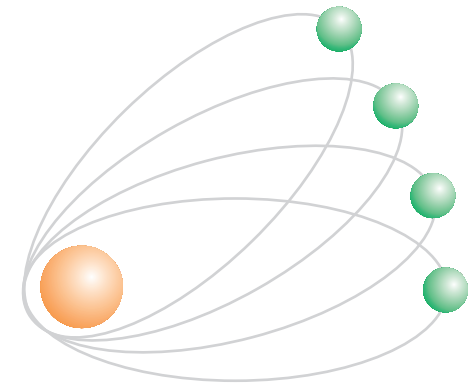
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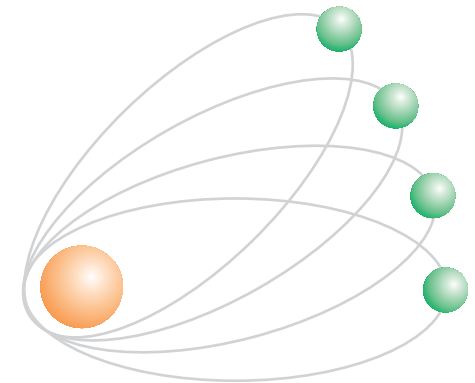
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 - interpreted curvature of spacetime as gravity
 - explained Mercury’s orbit without the “dark planet”



Testing general relativity

GR has many other testable predictions

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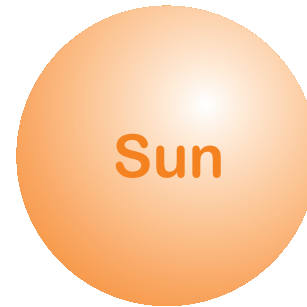
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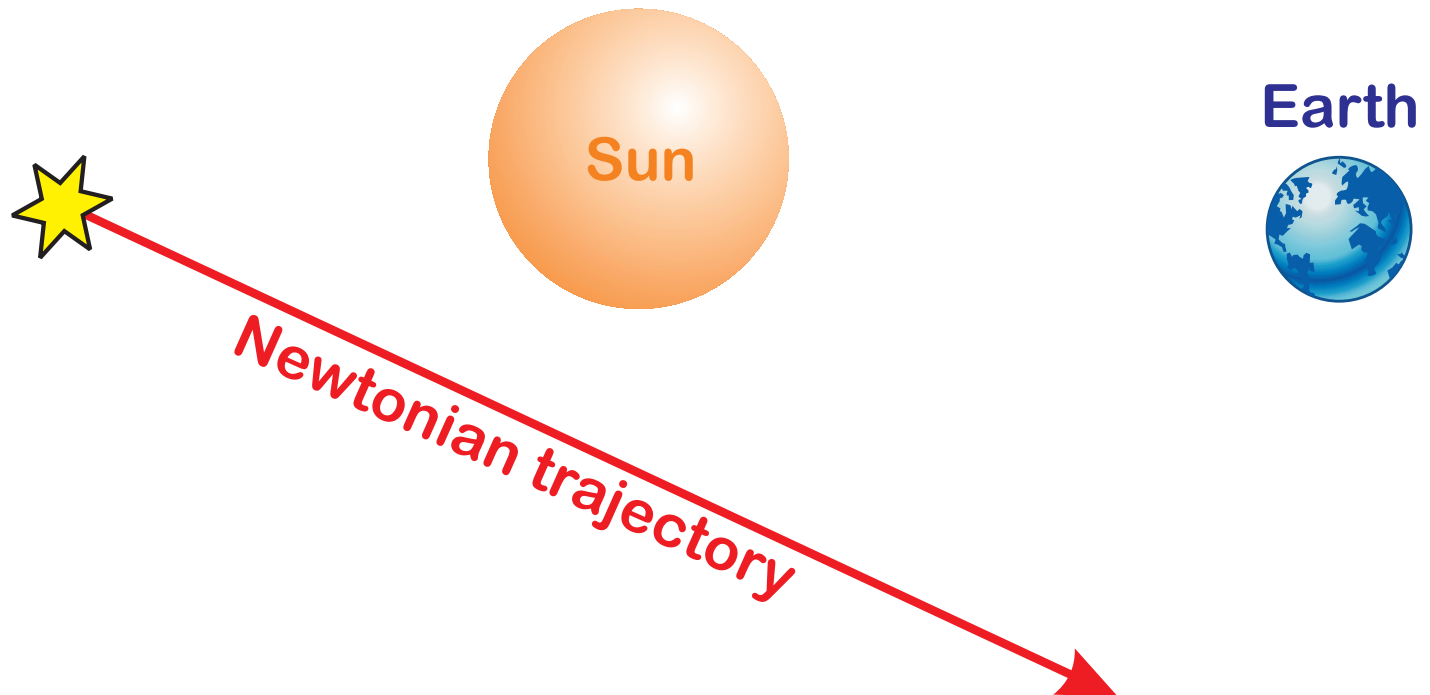
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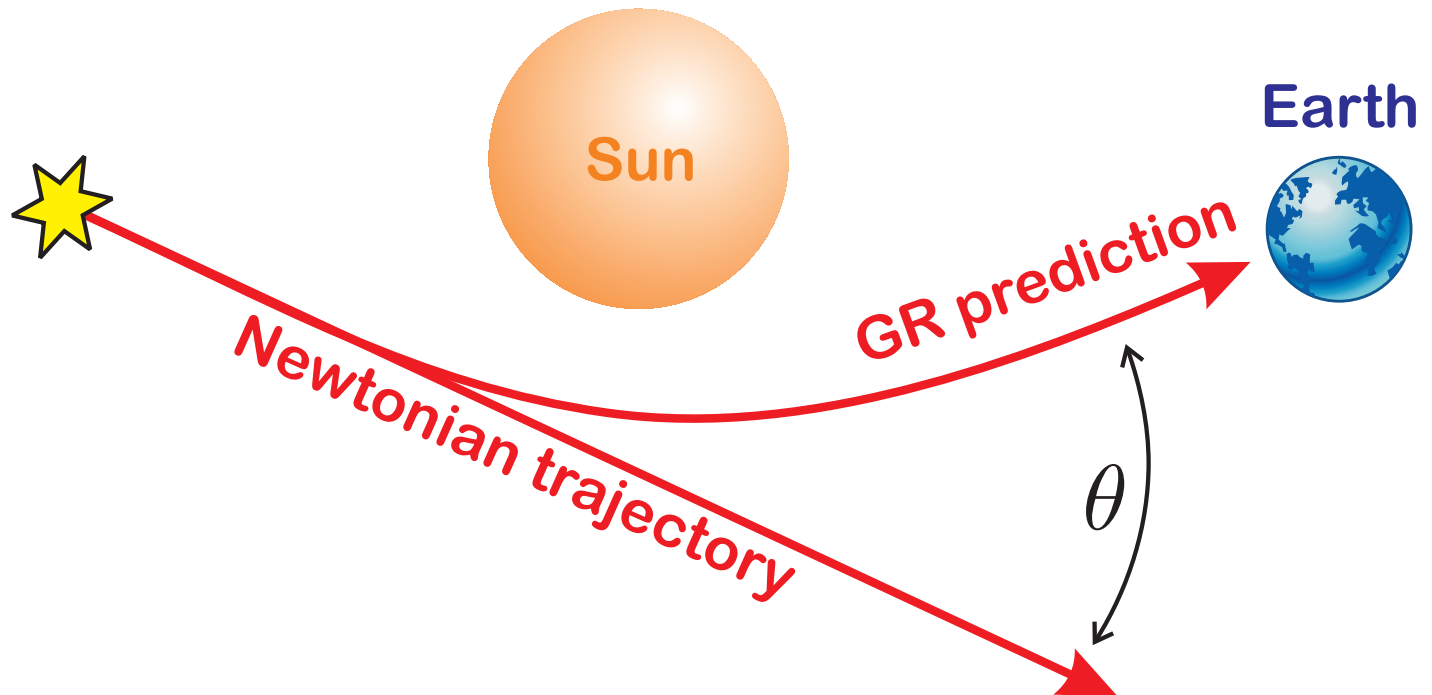
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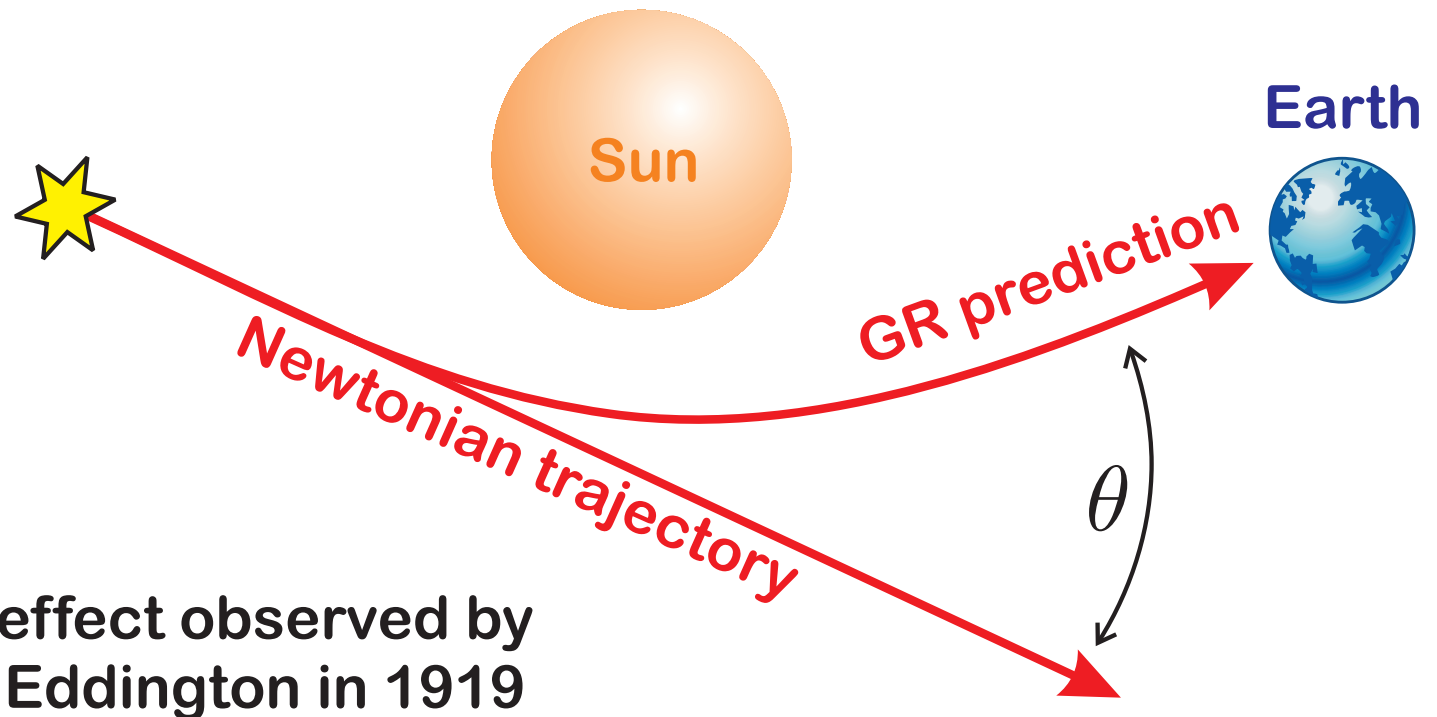
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effect observed by Eddington in 1919

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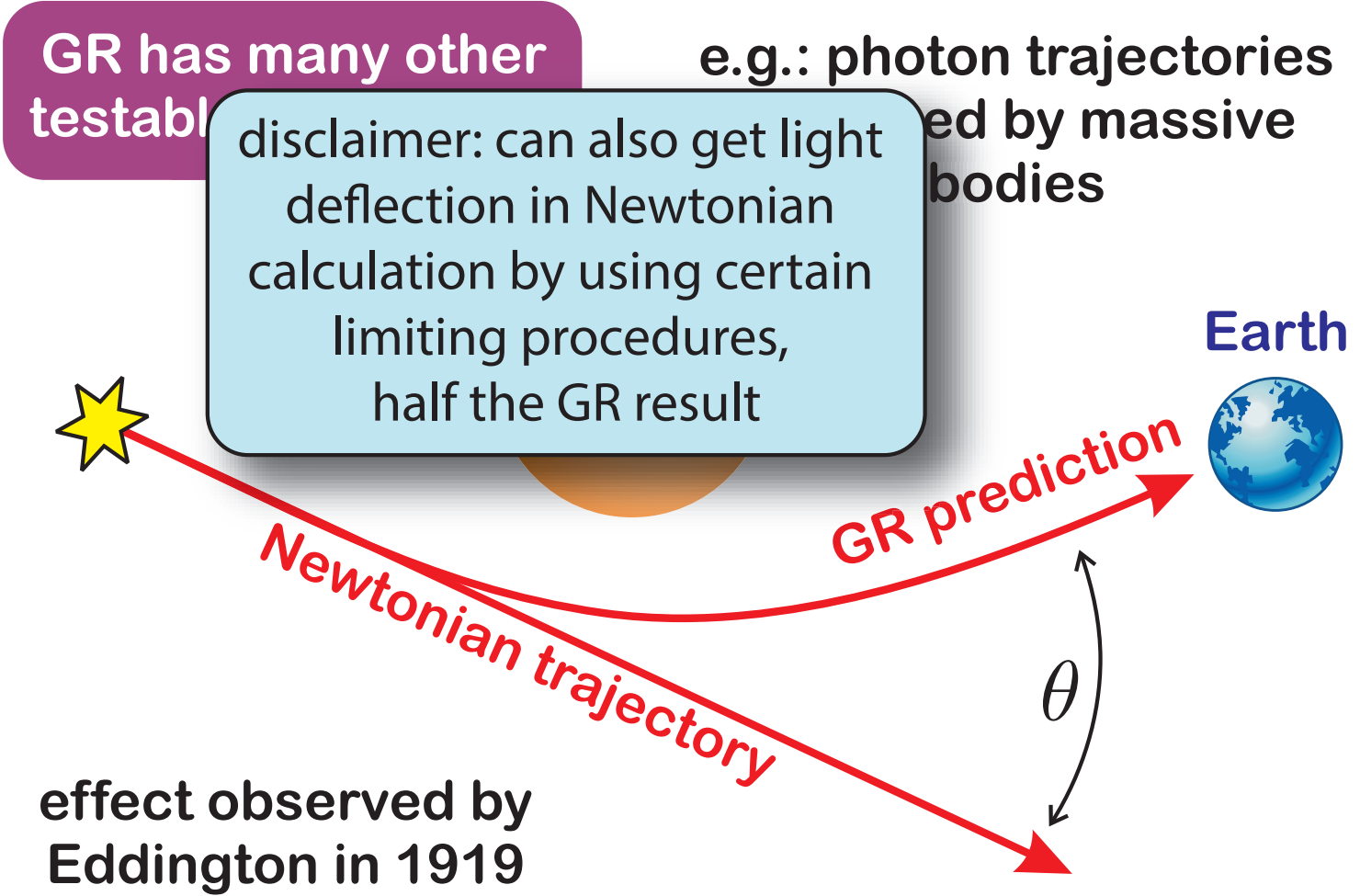
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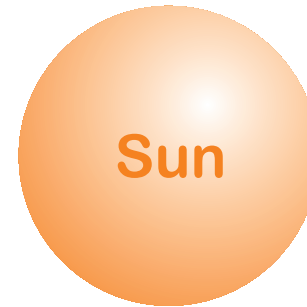
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Cassini spacecraft



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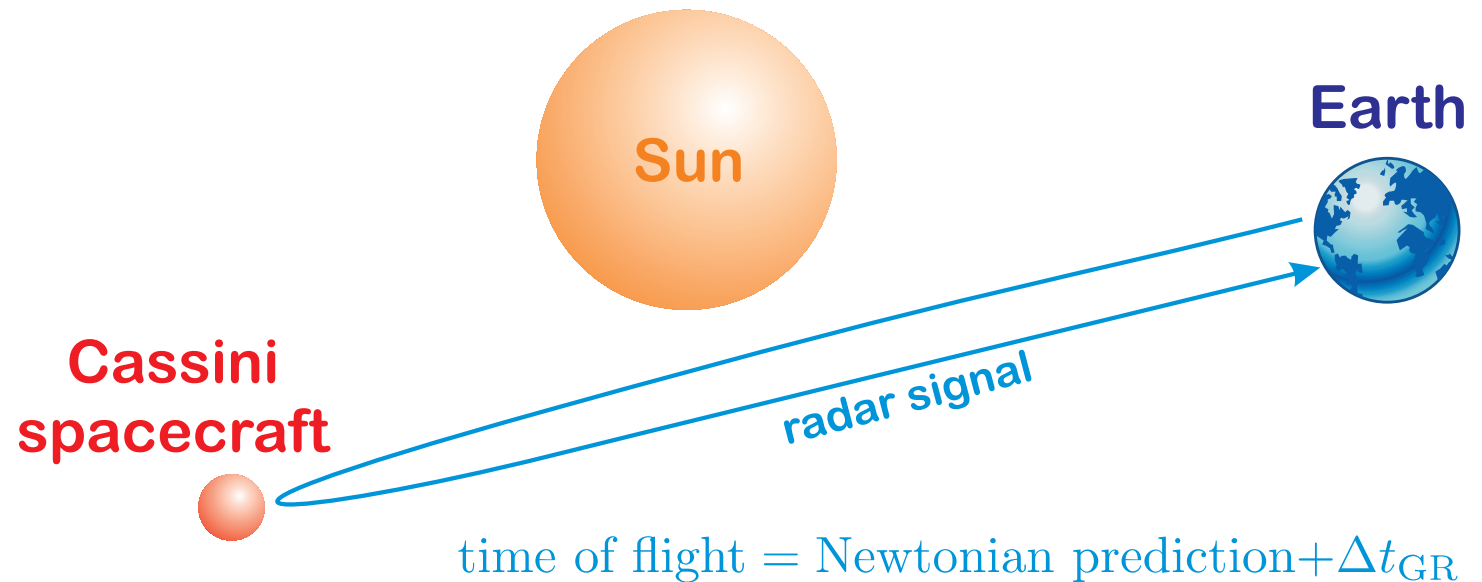
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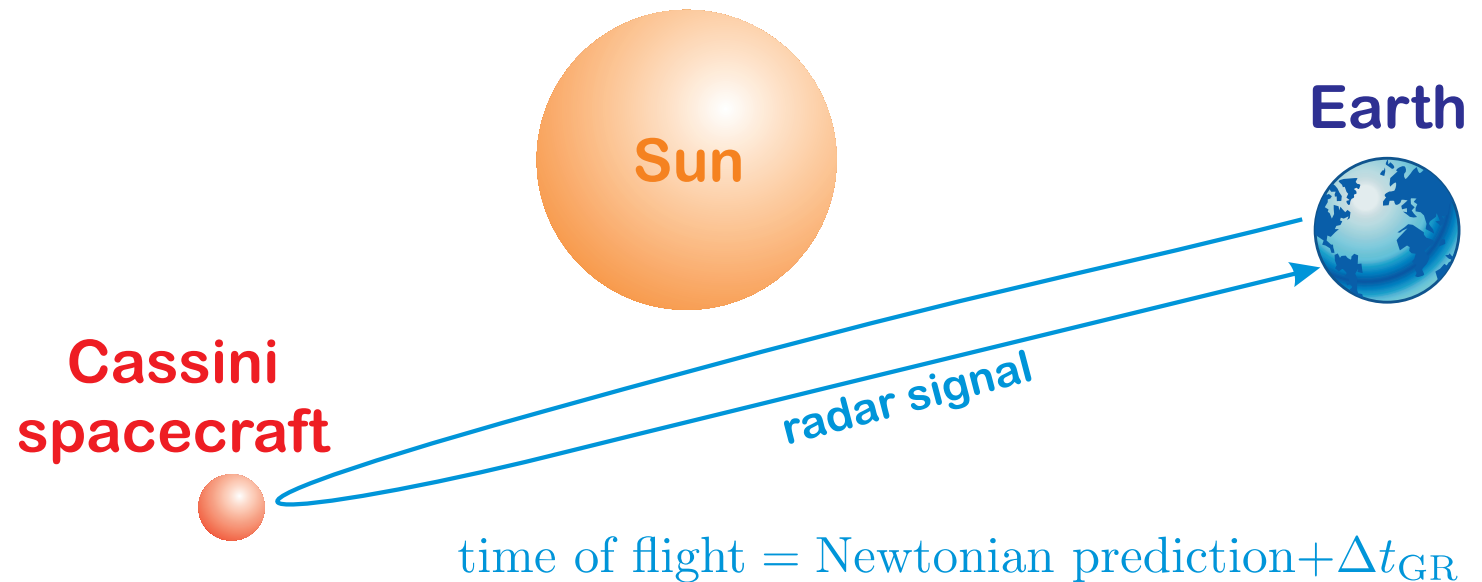
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measured Δt matches GR prediction to accuracy 10^{-5}

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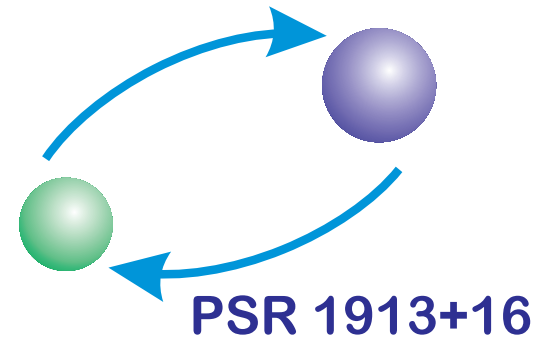
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e.g.: binary pulsars emit gravitational waves (GWs)

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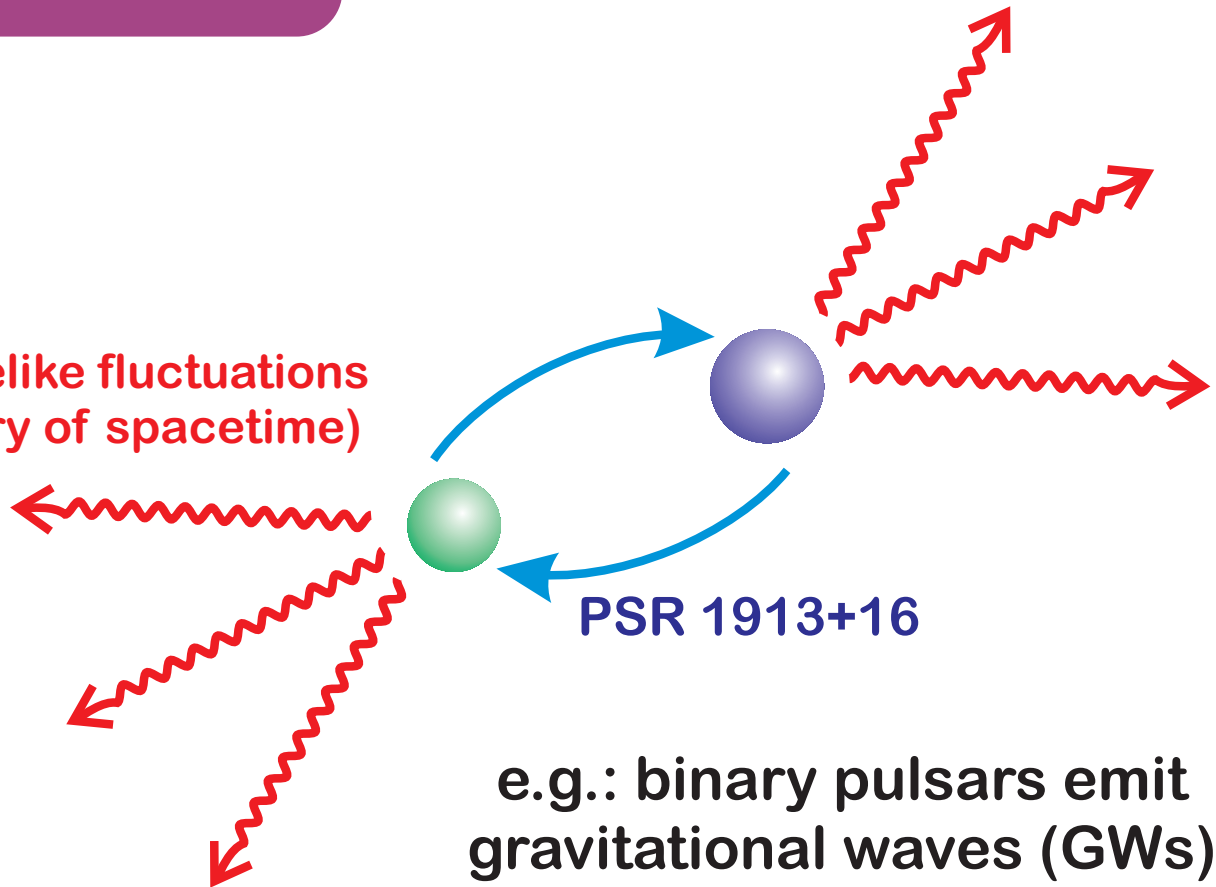
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GWs (wavelike fluctuations in geometry of spacetime)



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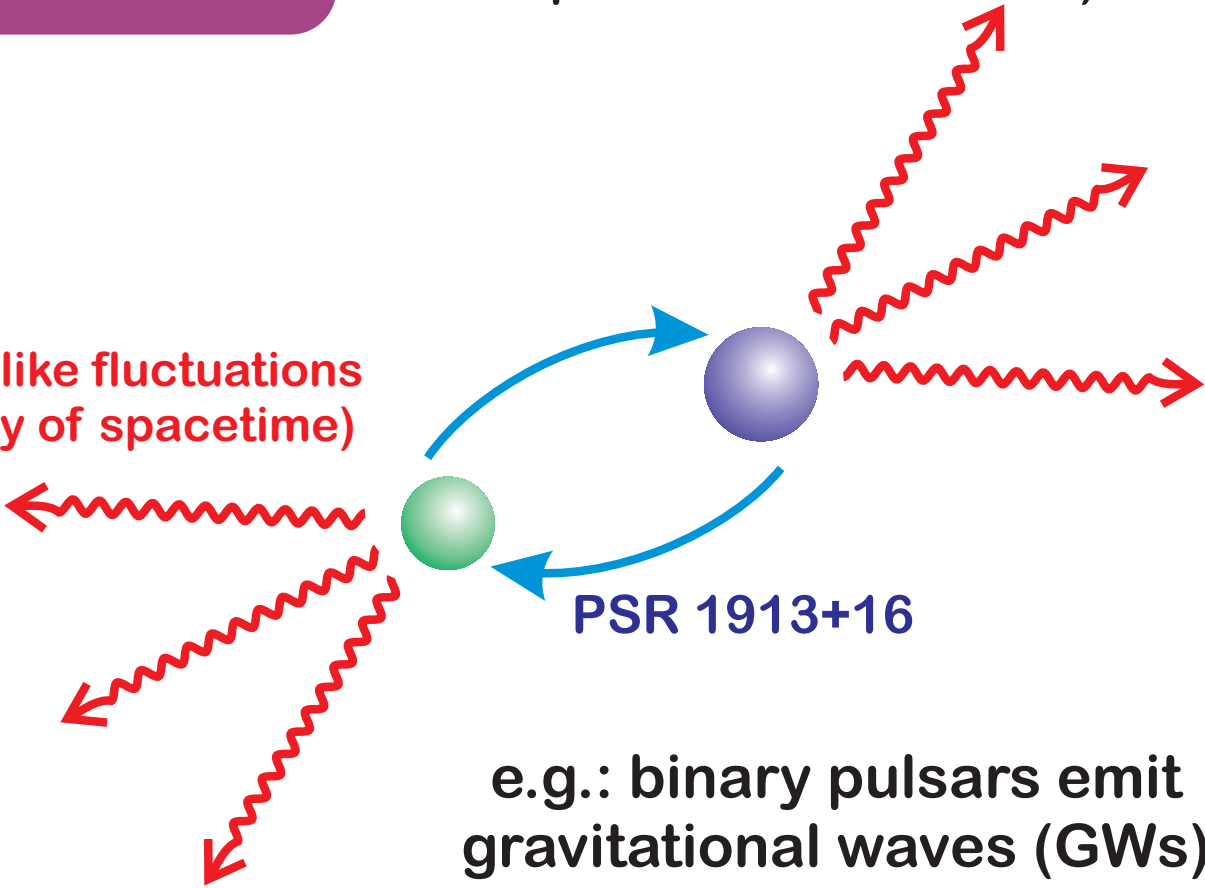
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(same effect as in Maxwell's theory: accelerating charges produce EM radiation)

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PSR 1913+16

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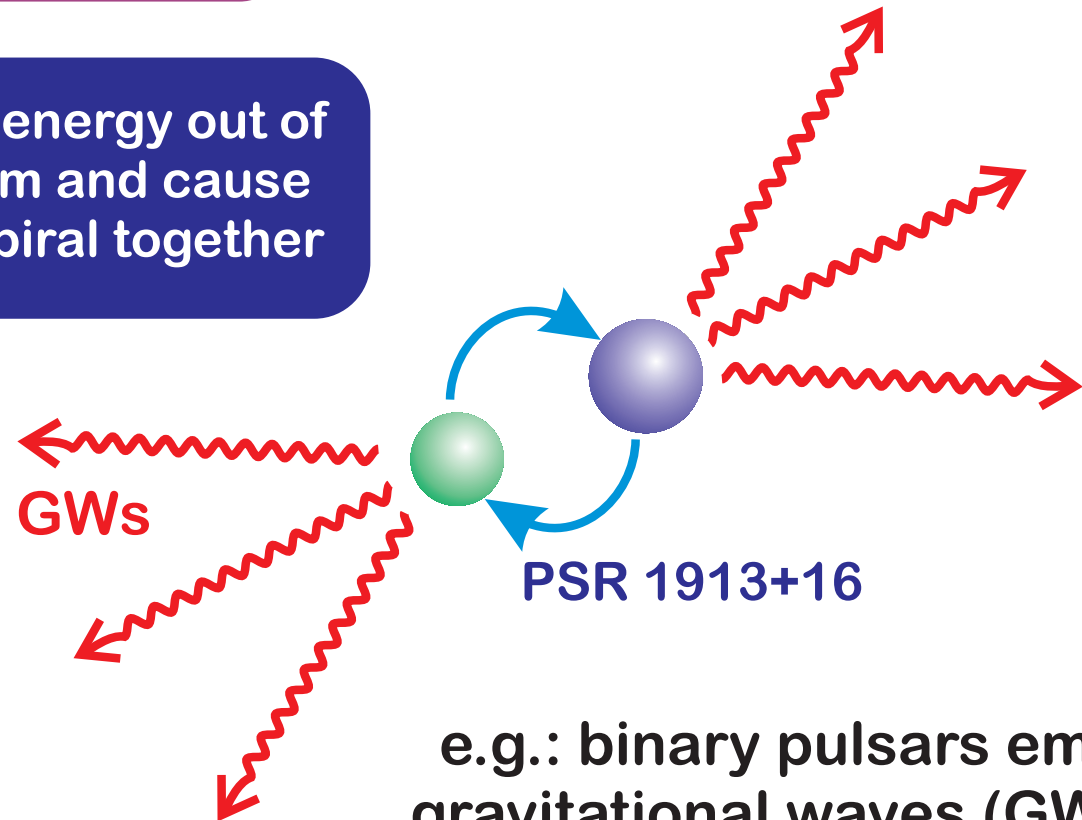
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GWs take energy out of the system and cause stars to spiral together

(same effect as in Maxwell's theory: accelerating charges produce EM radiation)



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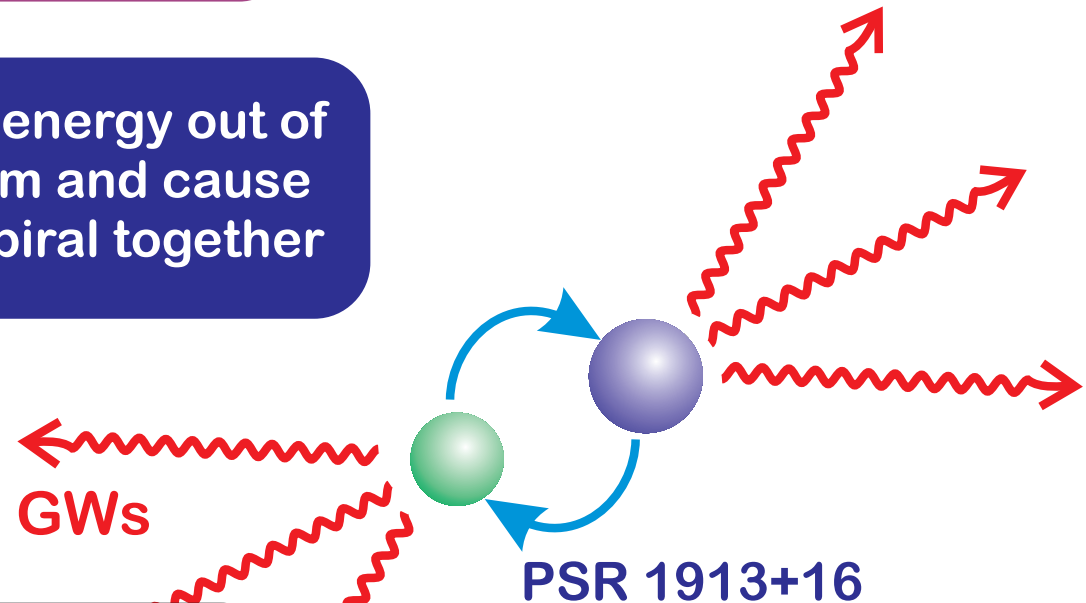
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GWs take energy out of the system and cause stars to spiral together

measured rate of inspiral matches GR prediction to accuracy 10^{-3}

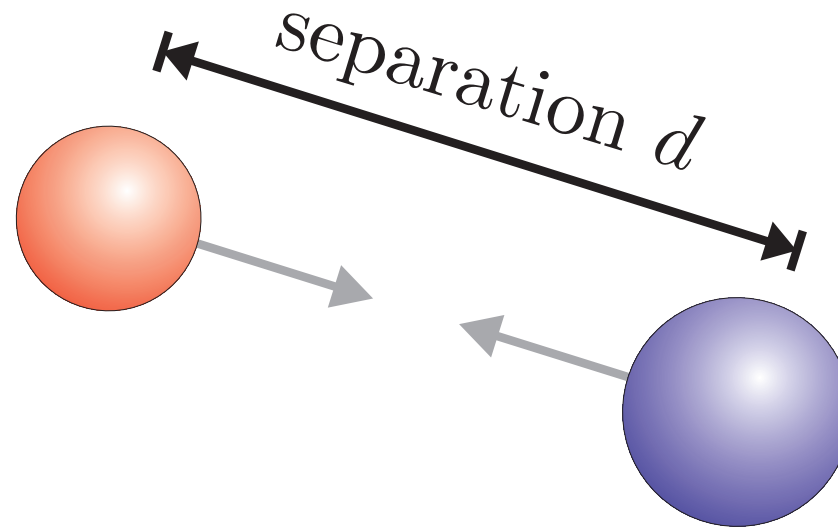
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can also test GR in the lab by measuring gravitational attraction directly

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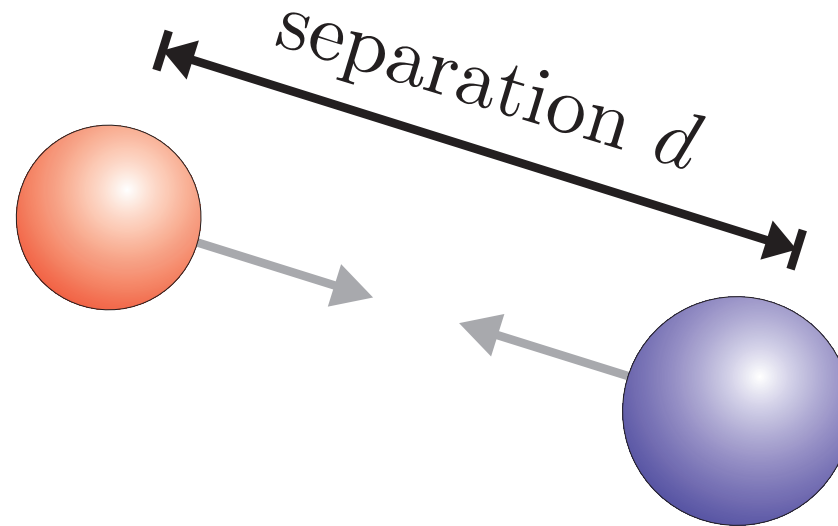
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GR prediction confirmed
for $d \gtrsim 50 \mu\text{m}$



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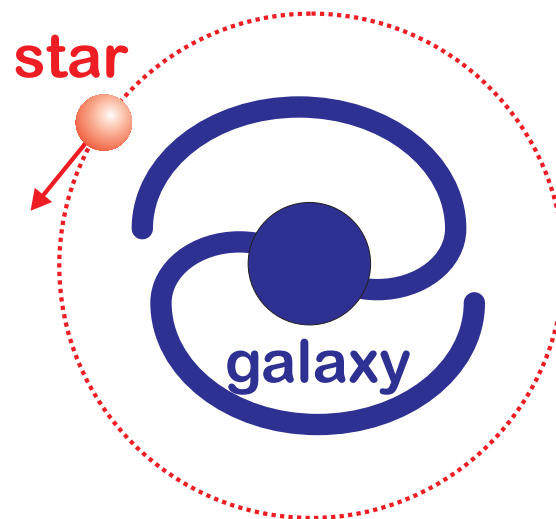
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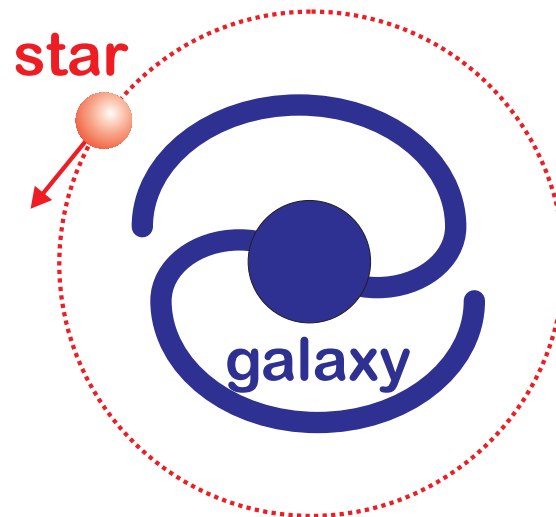
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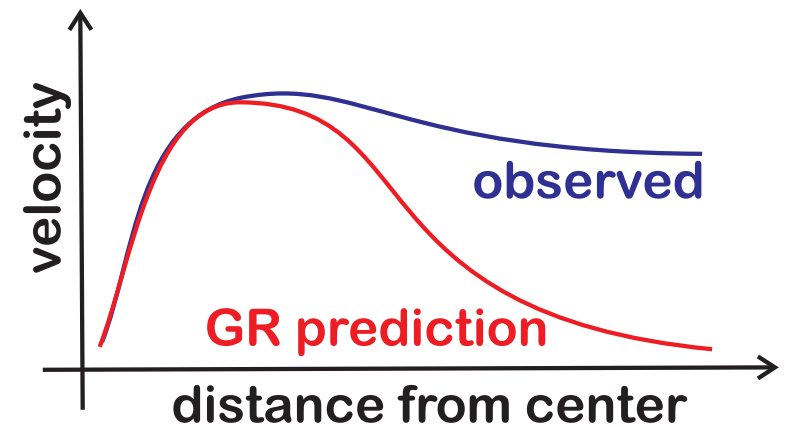
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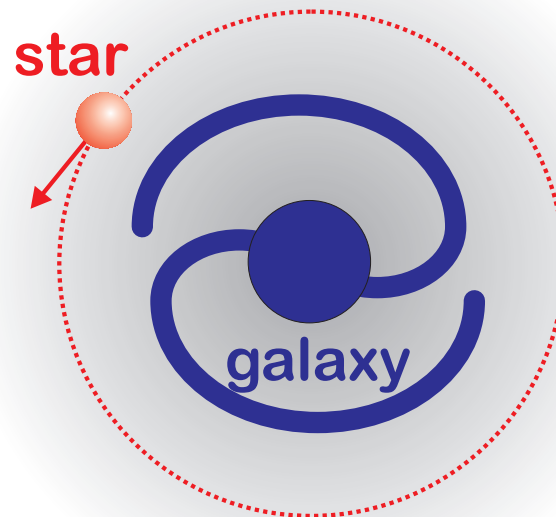
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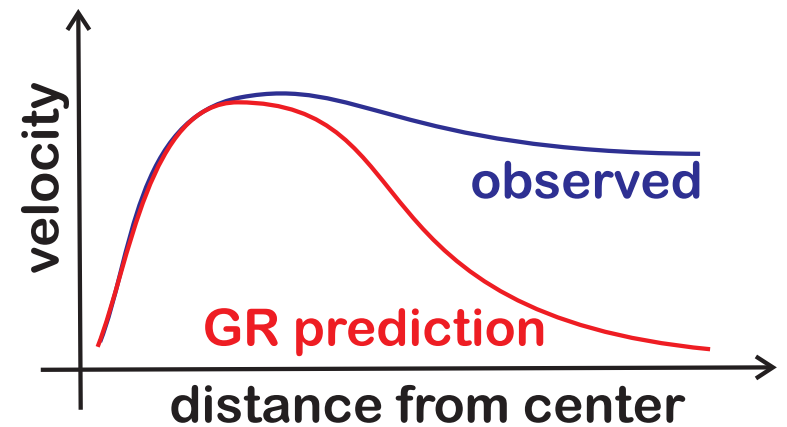
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discrepancy usually explained by the existence of "dark matter" haloes

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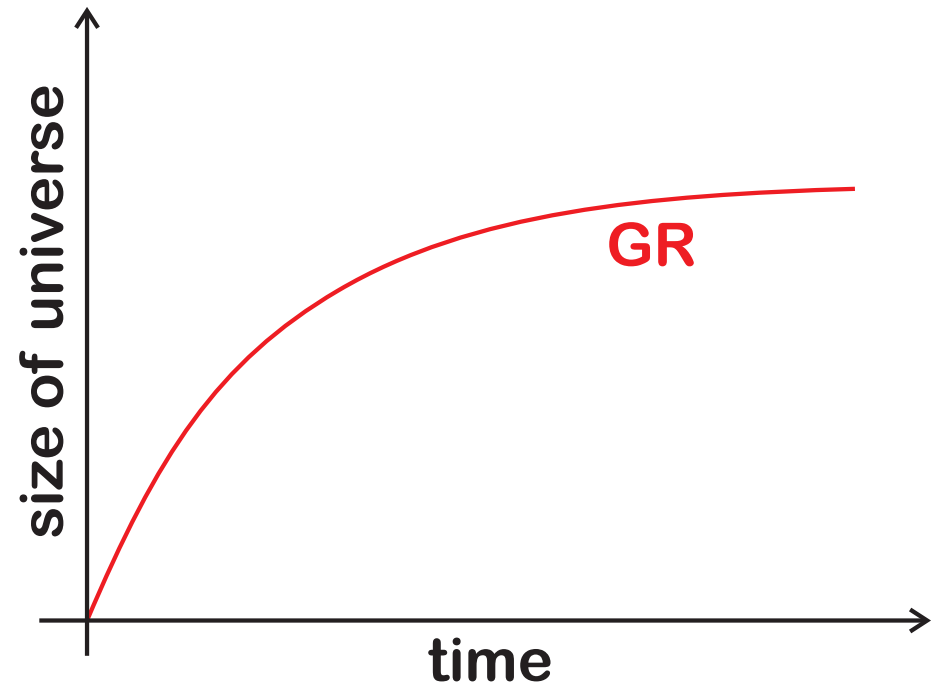
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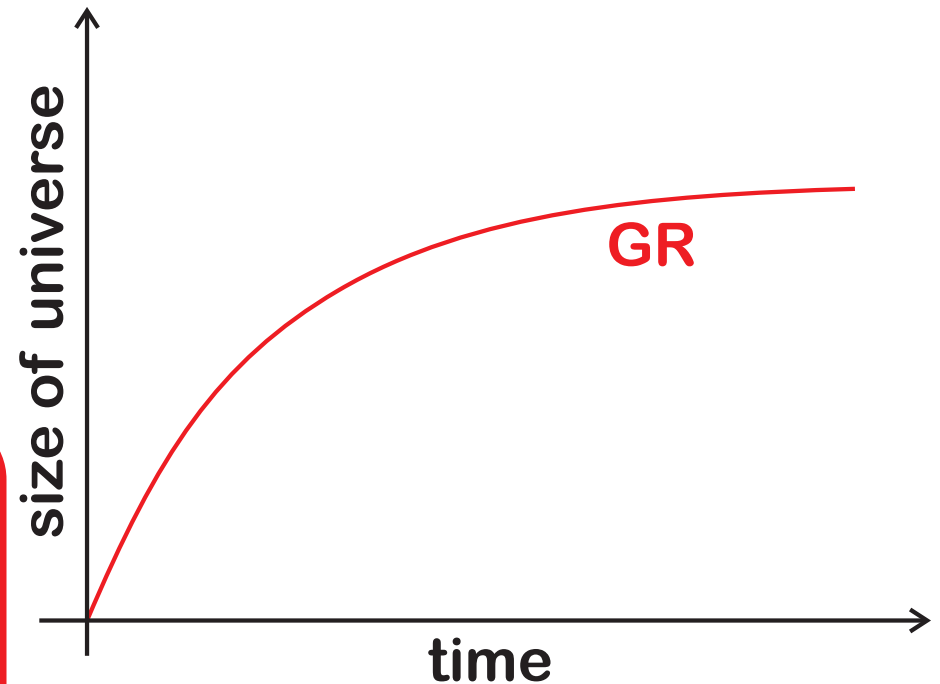
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in GR, gravity is an attractive force: so expansion should decelerate

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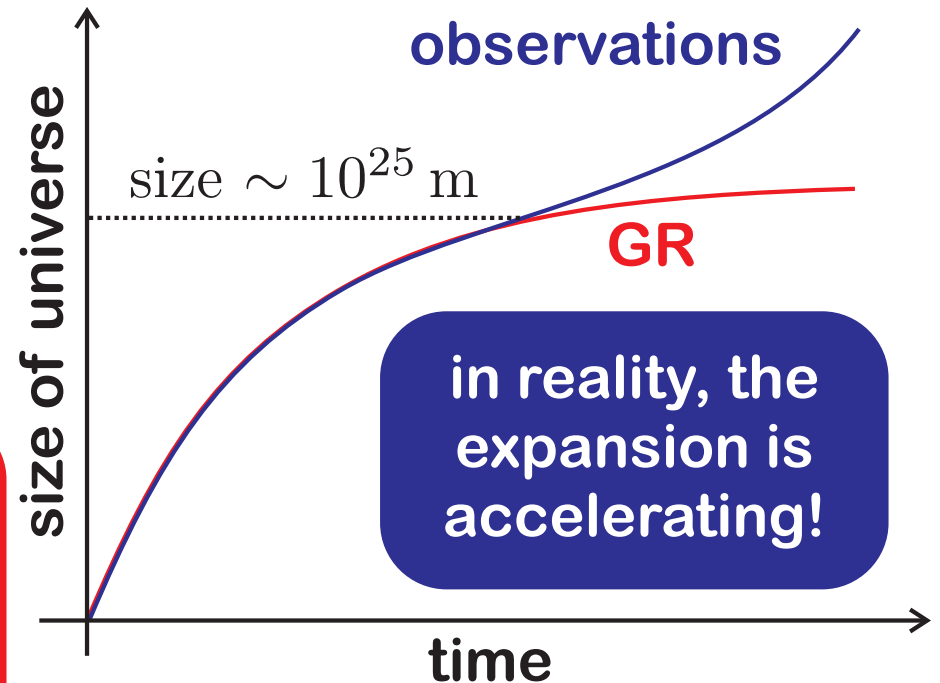
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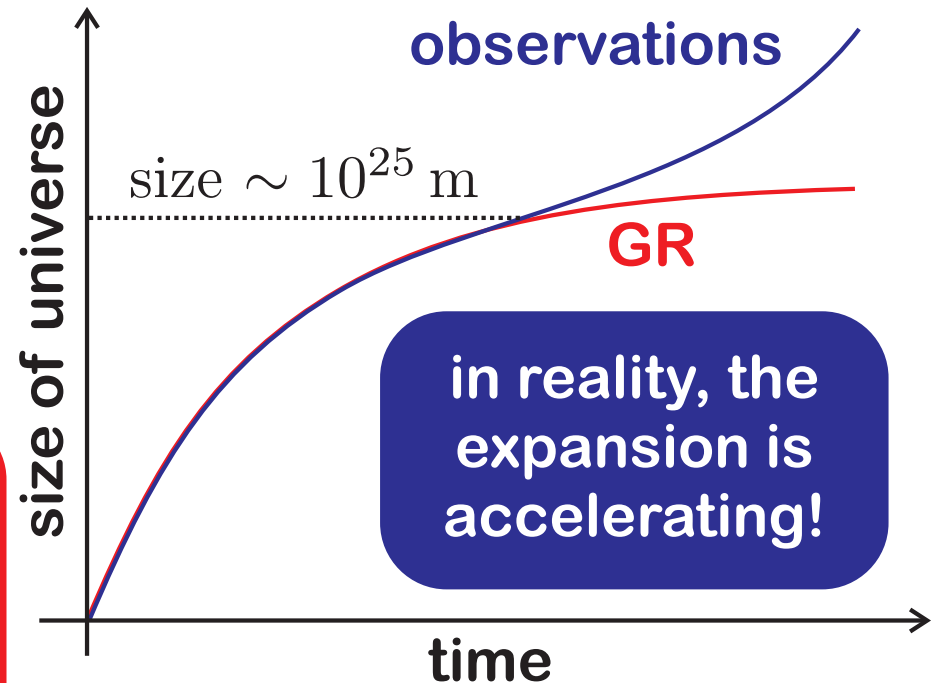
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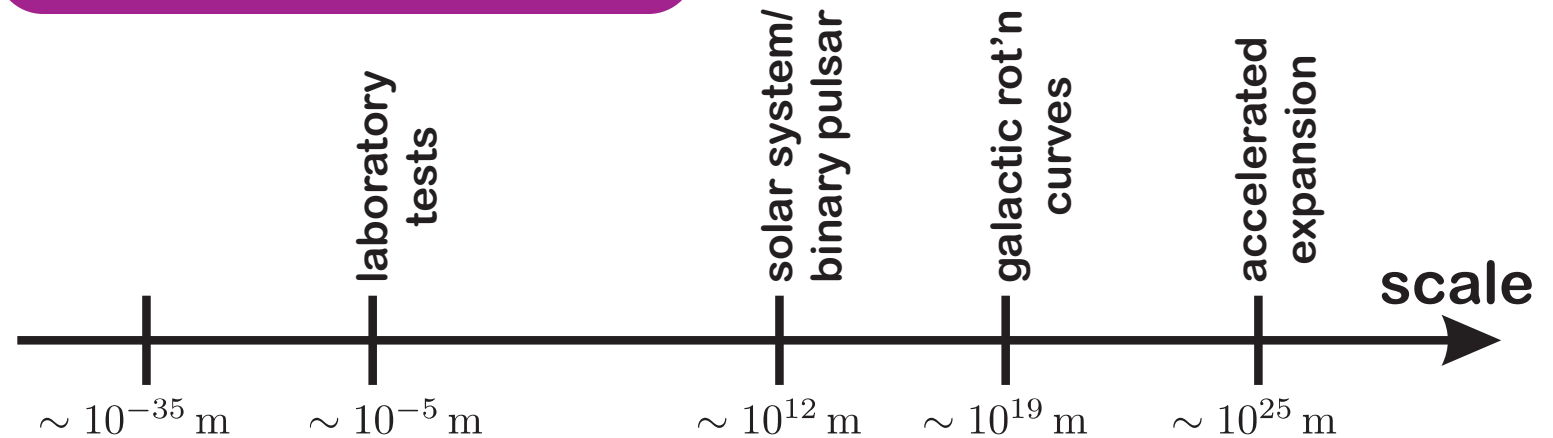
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different effects probe gravity at different scales



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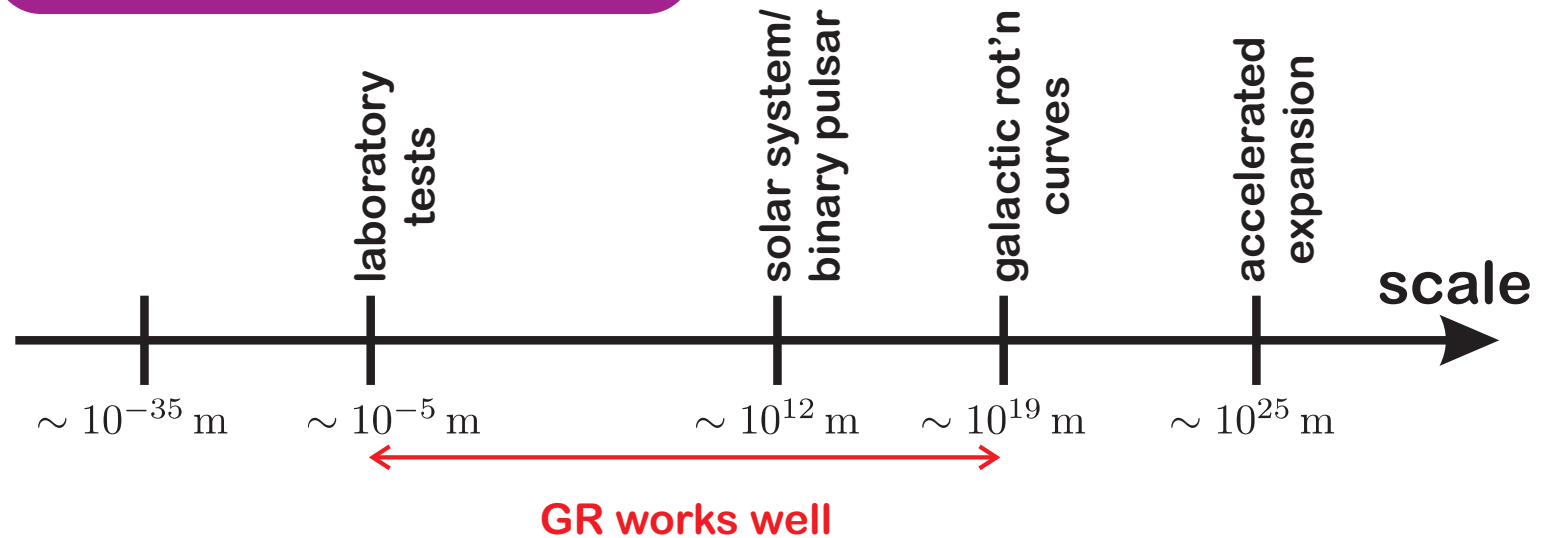
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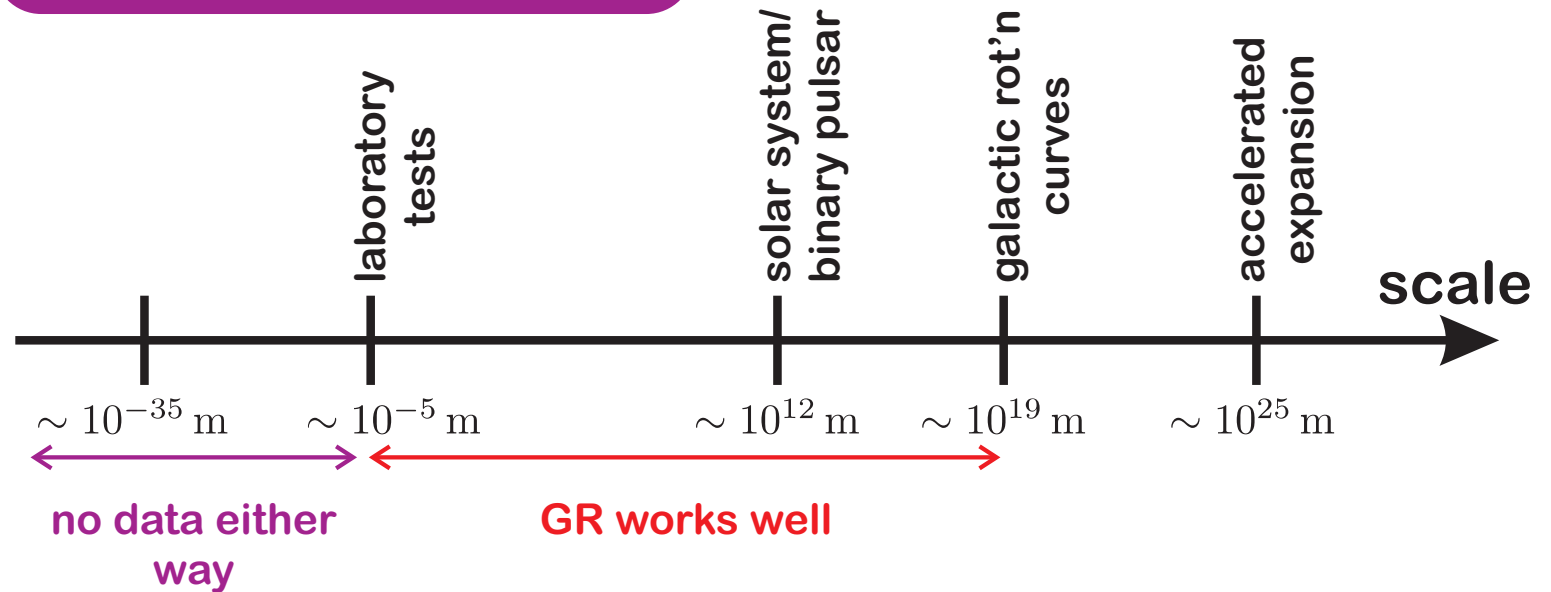
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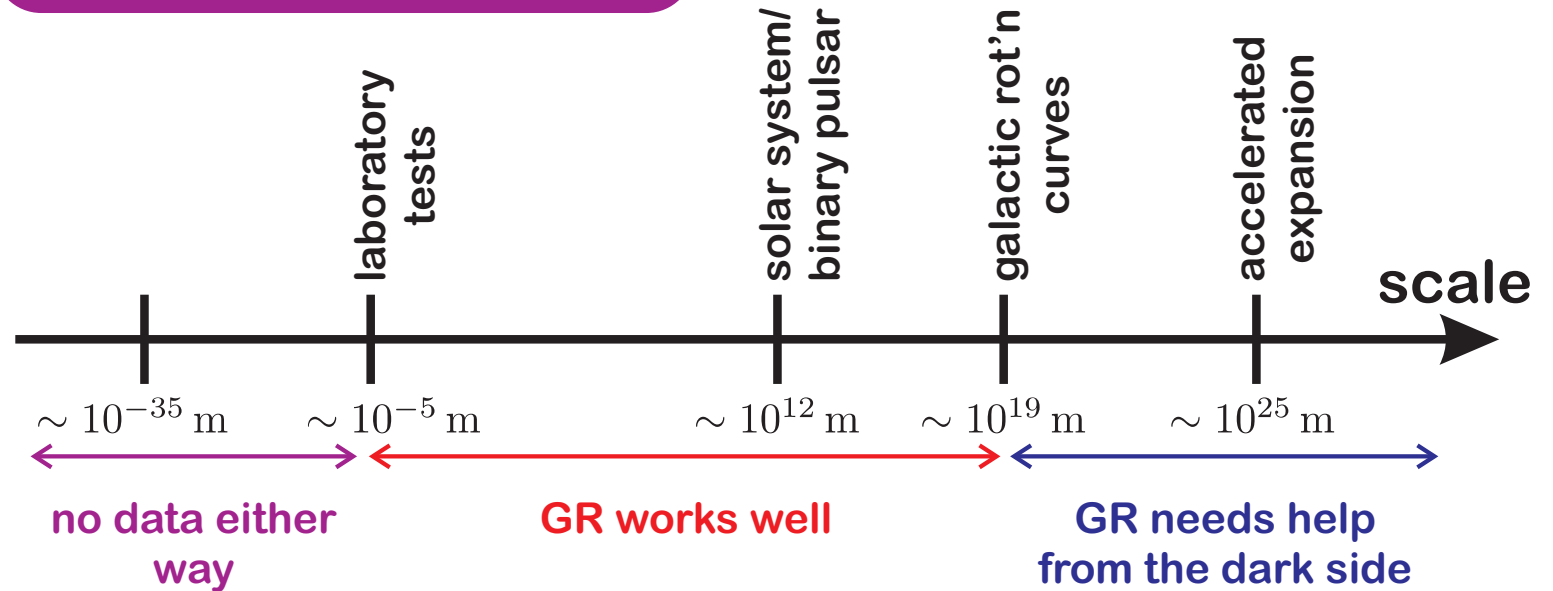
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History of gravity

Testing general relativity

Shortcomings of GR

How not to quantize gravity

Is this a problem?

“Observable” quantum gravity?

Hard to see

Inflation as a Planck scale microscope

Only testable QG prediction?

Small scale patches

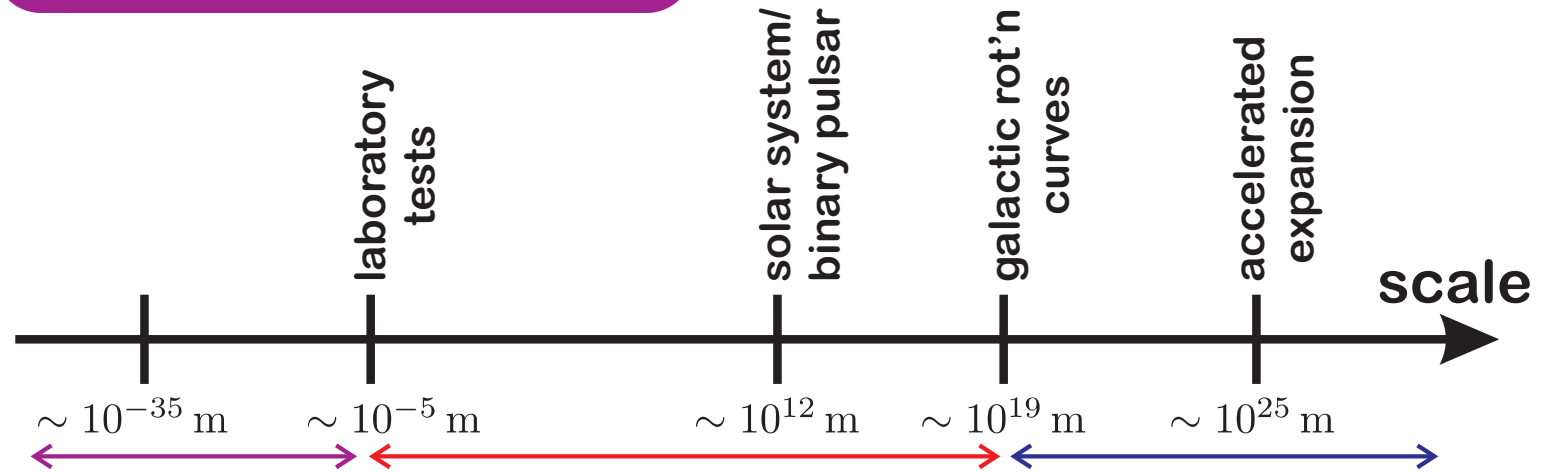
Large scale fixes

Where do we stand?

Shortcomings of GR

however, GR doesn't explain everything...

different effects probe gravity at different scales



no data either way

GR works well

GR needs help from the dark side

...or GR is not the correct theory of gravity on large scales



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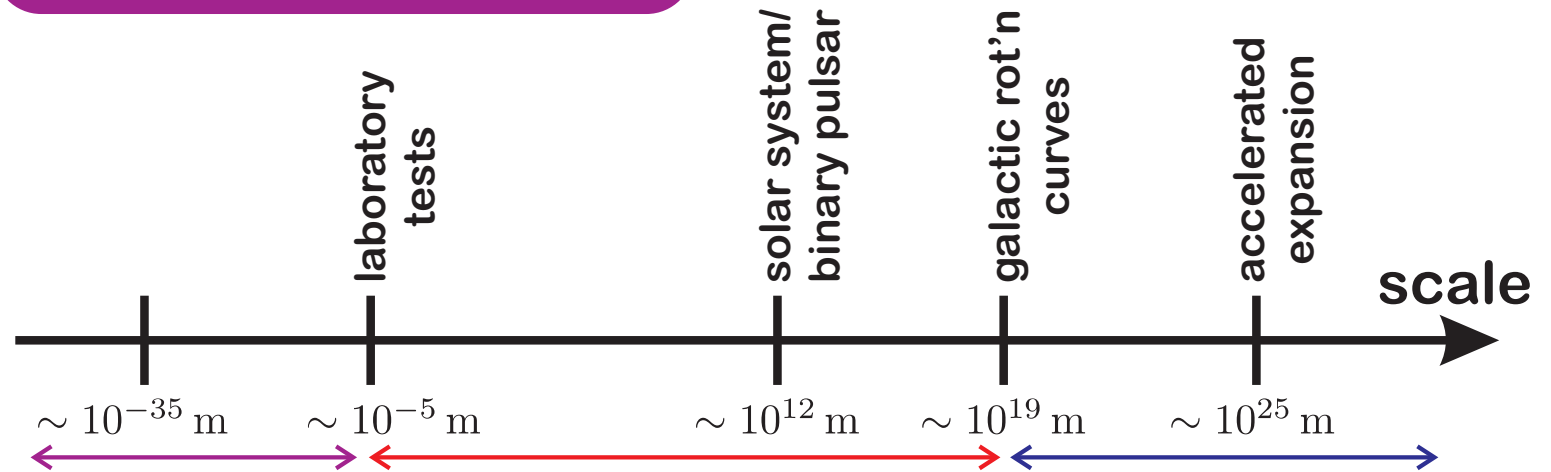
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is GR the right theory on small scales?

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 - that is, perturbative quantum gravity is non-renormalizable

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 - is there any experimental information available?

“Observable” quantum gravity?

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Where do we stand?

- how big are quantum gravity effects?
- the dimensionful constants in the theory:
 - Newton’s constant G
 - Planck’s constant \hbar
 - the speed of light c

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$$M_{\text{Pl}} = \sqrt{\frac{\hbar c}{G}} \sim \frac{10^{19} \text{ GeV}}{c^2} \quad \ell_{\text{Pl}} = \sqrt{\frac{\hbar G}{c^3}} \sim 10^{-35} \text{ m}$$

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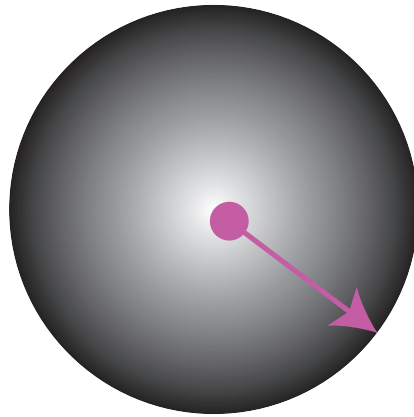
Only testable QG prediction?

Small scale patches

Large scale fixes

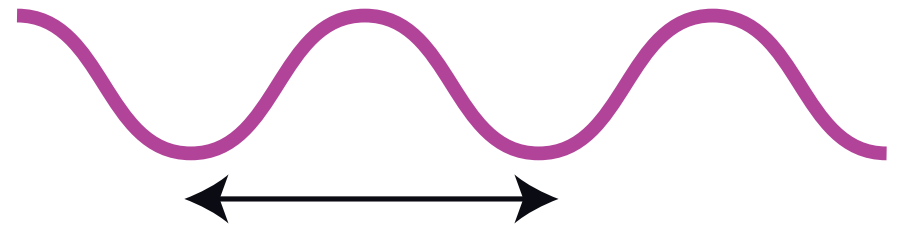
Where do we stand?

black hole of mass M



radius $\frac{2GM}{c^2}$

quantum particle of mass M



Compton wavelength $\frac{2\pi\hbar}{Mc}$

$$\text{set them equal: } M = \sqrt{\frac{\pi\hbar c}{G}} = \sqrt{\pi} M_{\text{Pl}}$$

a black hole with mass M_{Pl} has a radius \approx Compton wavelength

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- physics intuition: quantum gravity effects will become tangible for (individual particle) energies $\gtrsim M_{\text{Pl}}$ or distances $\lesssim \ell_{\text{Pl}}$

This means its not easy to observe quantum gravity effects:

phenomenon	typical size ℓ	ℓ_{Pl}/ℓ
solar mass black hole	10^3 m	10^{-38}
hydrogen atom	10^{-10} m	10^{-25}
proton	10^{-15} m	10^{-20}

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phenomenon	typical energy E	$E/M_{\text{Pl}}c^2$
hydrogen atom	10^{-9} GeV	10^{-28}
large hadron collider	10^3 GeV	10^{-16}
ultra high energy cosmic rays	10^9 GeV	10^{-8}
cosmological inflation	10^{16} GeV	10^{-3}

Successes and failures

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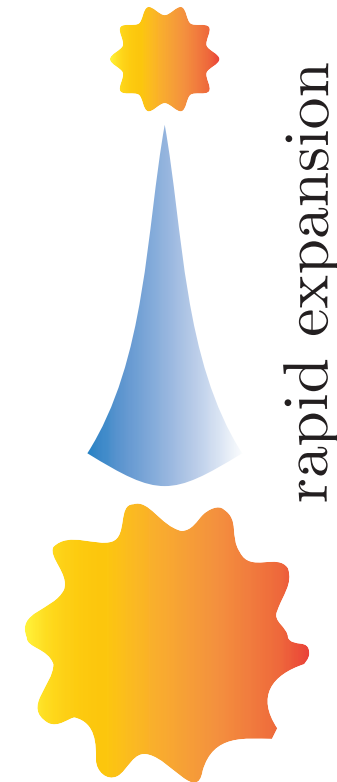
Large scale fixes

Where do we stand?

Inflation as a Planck scale microscope

inflation provides a mechanism for the generation of structure in the early universe

quantum fluctuations when universe is small



macroscopic inhomogeneities

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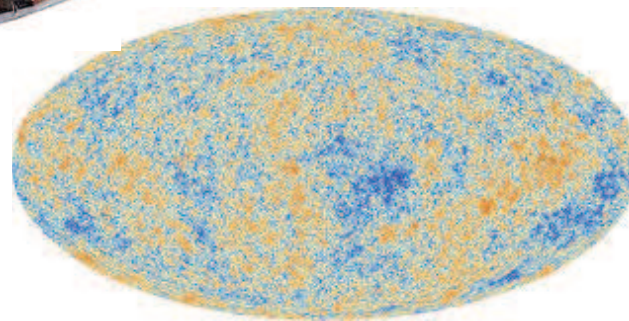
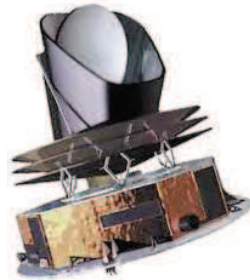
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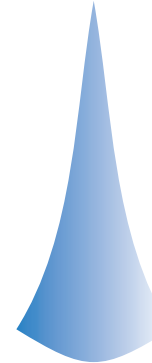
quantum fluctuations when universe is small

leads to predictions for the cosmic microwave background

Planck satellite



temperature anisotropies



rapid expansion



macroscopic inhomogeneities

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these have wavelength $\ll \ell_{\text{Pl}}$ and should be sensitive to quantum gravity

quantum fluctuations when universe is small



rapid expansion

Kempf: “A Planck scale microscope”

these are big enough to see in the distribution of objects in the sky



macroscopic inhomogeneities

\Rightarrow we could see quantum gravity effects in large scale structure of universe

Primordial GWs: testable prediction?

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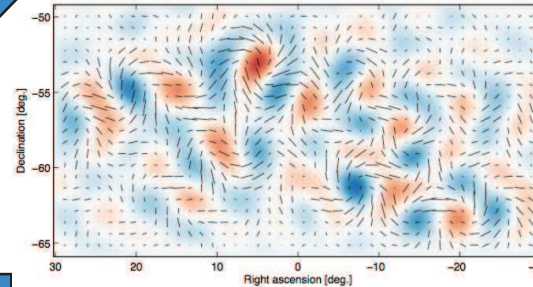
Large scale fixes

Where do we stand?

tensor modes: purely gravitational fluctuations enhanced during inflation



give rise to characteristic polarization in CMB

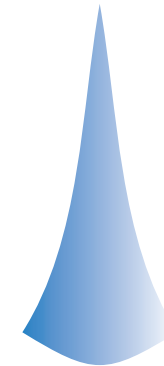


observation would provide circumstantial evidence for quantized linear gravitation perturbations



finiteness would suggest that gravity ought to be quantized and renormalizable

quantum fluctuations when universe is small



rapid expansion

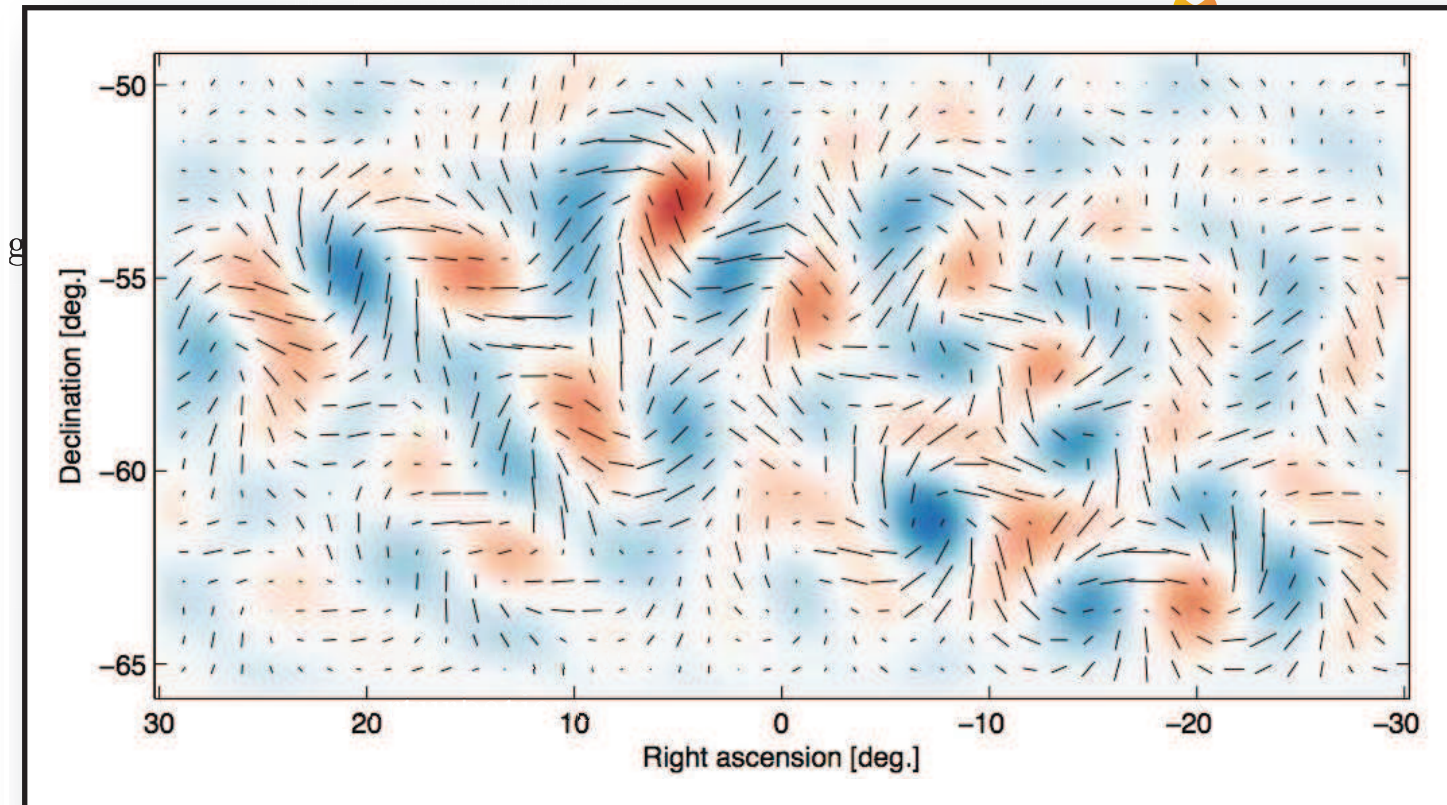


macroscopic inhomogeneities

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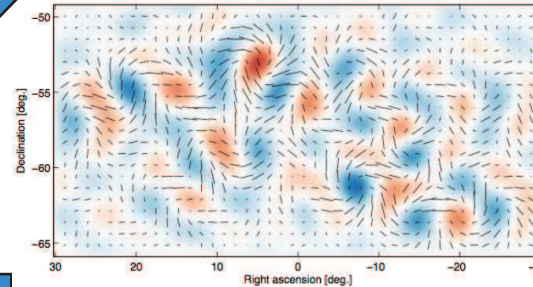
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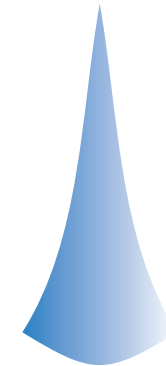


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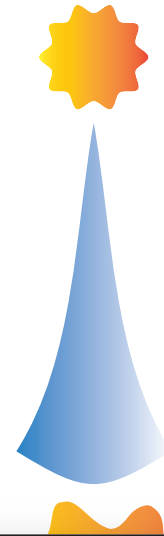
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give
p

quantum fluctuations
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**DETECTION
ANNOUNCED MONDAY,
MARCH 17 BY BICEP2!**

macroscopic
inhomogenieties

finiteness would suggest that
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Successes and failures

Small scale patches

Solution 1: quantizing
the wrong theory

Solution 2: right theory
but wrong quantization

... and the answer is

Large scale fixes

Where do we stand?

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Solution 1: quantizing the wrong theory

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- if GR is cannot be consistently quantized, maybe it is not the correct theory of gravity

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Successes and failures

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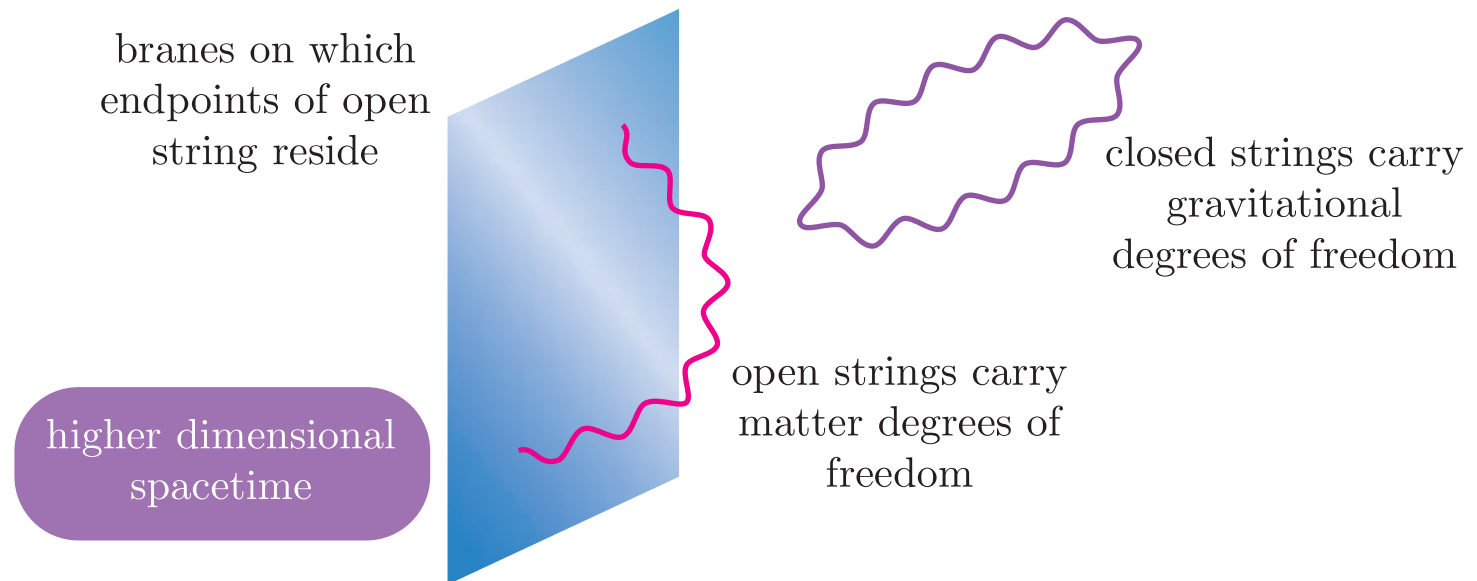
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Large scale fixes

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 - **string theory:** gravitational field is actually induced by vibrations of higher dimensional objects (strings and branes)



Solution 1: quantizing the wrong theory

Successes and failures

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 - **higher curvature and Horava-Lifshitz theories:** GR is effective version of a more complicated (4D) theory that can be consistently quantized
 - **pros:** the theories are selected to avoid divergences when quantized

Solution 2: right theory but wrong quantization

Successes and failures

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Where do we stand?

- if we try to quantize gravitational perturbations and get inconsistent results, maybe we ought to try something else?

Solution 2: right theory but wrong quantization

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Solution 1: quantizing the wrong theory

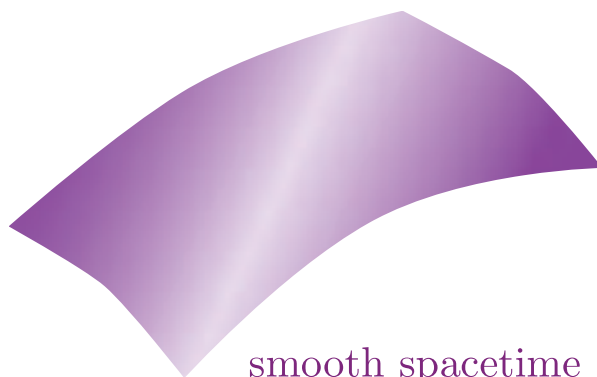
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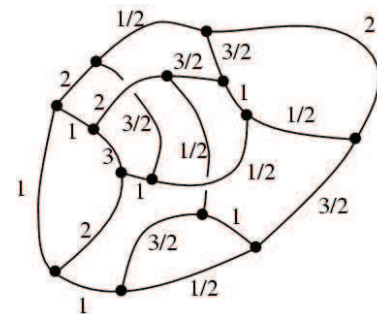
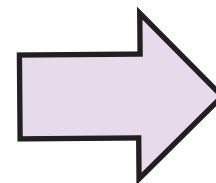
Large scale fixes

Where do we stand?

- if we try to quantize gravitational perturbations and get inconsistent results, maybe we ought to try something else?
- **loop quantum gravity:** tries to quantize GR non-perturbatively



smooth spacetime geometry



connected graph labelled by spin indices

Solution 2: right theory but wrong quantization

Successes and failures

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Solution 1: quantizing the wrong theory

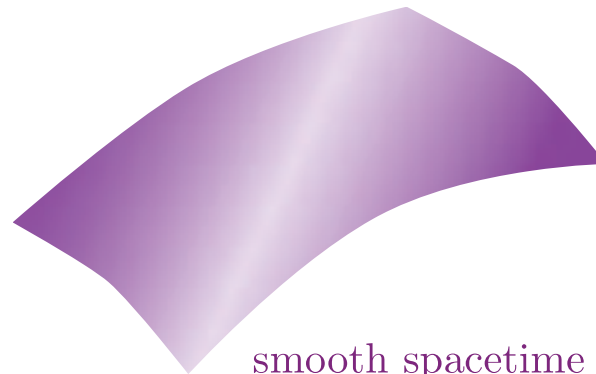
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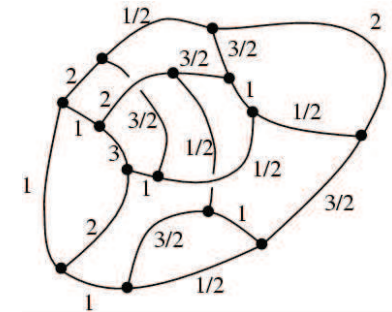
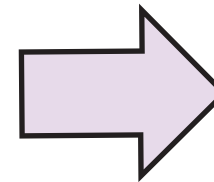
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smooth spacetime geometry



connected graph labelled by spin indices

- **pros:** no exotic classical theory of gravity; succeeds because of choice of variables and non-standard polymer quantization

...and the answer is

Successes and failures

Small scale patches

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Where do we stand?



...and the answer is

- in the absence of experimental facts, the only way to distinguish models is consistency checks/theoretical bias

Successes and failures

Small scale patches

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Large scale fixes

Where do we stand?

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Successes and failures

Small scale patches

Solution 1: quantizing
the wrong theory

Solution 2: right theory
but wrong quantization

... and the answer is

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- in the absence of experimental facts, the only way to distinguish models is consistency checks/theoretical bias
- there are criticisms of each approach:

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Where do we stand?

- in the absence of experimental facts, the only way to distinguish models is consistency checks/theoretical bias
- there are criticisms of each approach:
 - string theory involves a background higher dimensional spacetime; i.e. not really a quantization of geometry
 - really hard to calculate things in loop quantum gravity; not even sure how classical GR is recovered

Successes and failures

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Large scale fixes

The dark matter
problem

The dark energy
problem

Other ideas

Where do we stand?

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The dark matter problem

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Where do we stand?

- do we need to modify gravity to account for dark matter?

The dark matter problem

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Where do we stand?

- do we need to modify gravity to account for dark matter?
- probably not . . . lots of evidence for dark matter other than galactic rotation curves

The dark matter problem

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Large scale fixes

The dark matter problem

The dark energy problem

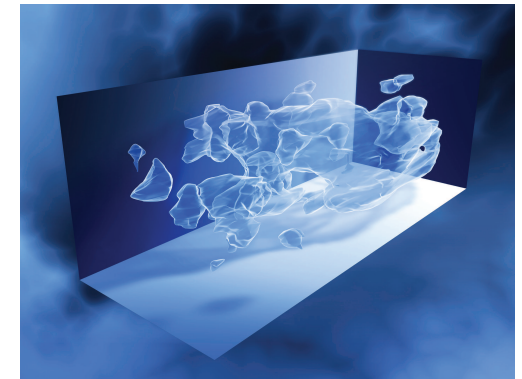
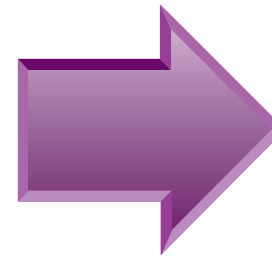
Other ideas

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 - for example: weak gravitational lensing lets us map dark matter distribution in galactic clusters



distortions in galactic shapes



map of dark matter distribution

The dark matter problem

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 - **indirect searches:** observe by-products of dark matter annihilations in the Milky Way

The dark energy problem

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Other ideas

Where do we stand?

- problem of explaining the late time acceleration of universe one of the biggest in physics

The dark energy problem

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The dark energy
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Other ideas

Where do we stand?

- problem of explaining the late time acceleration of universe one of the biggest in physics
- if it has a resolution like dark matter, the matter is very strange indeed

The dark energy problem

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 - observational value of Λ is ridiculously small, hard to “explain” it from other theories with dimensionful constants

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dark
energy

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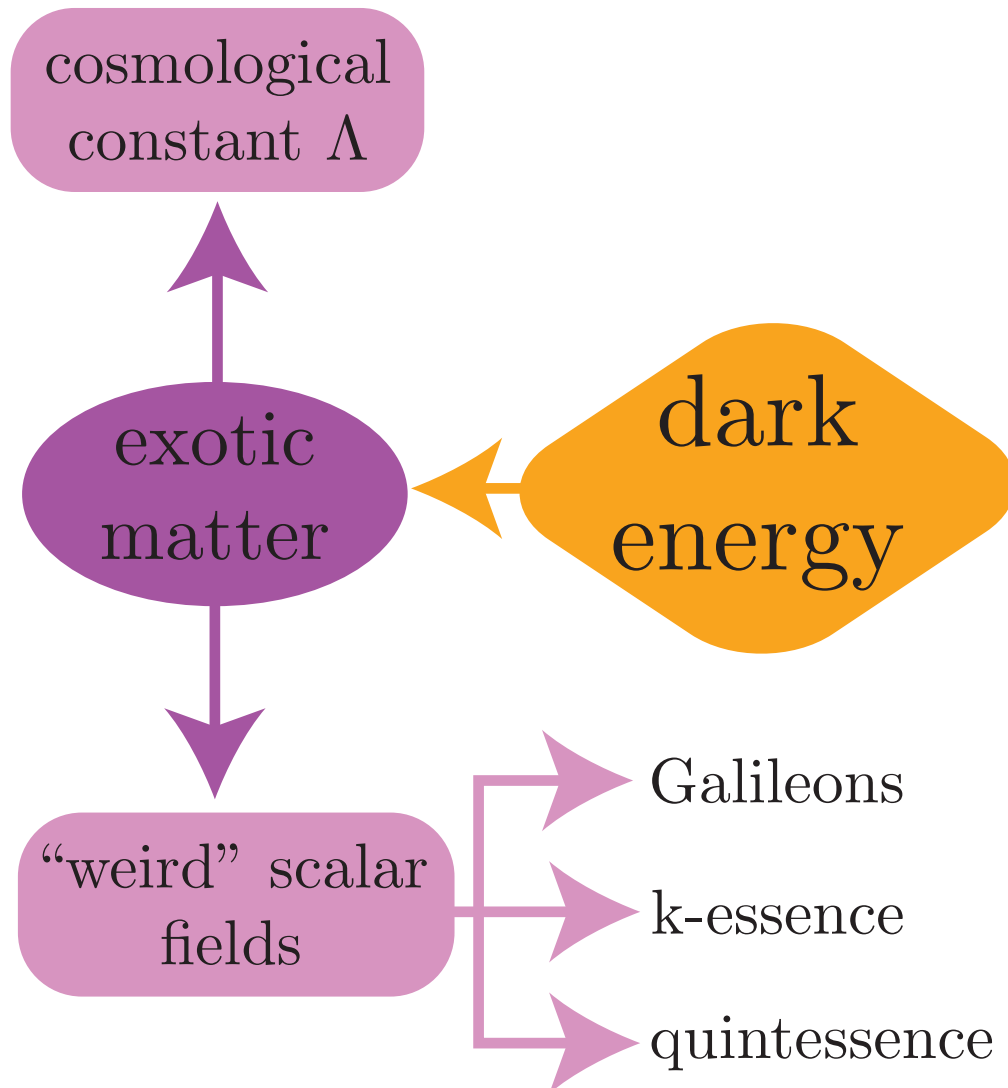
Large scale fixes

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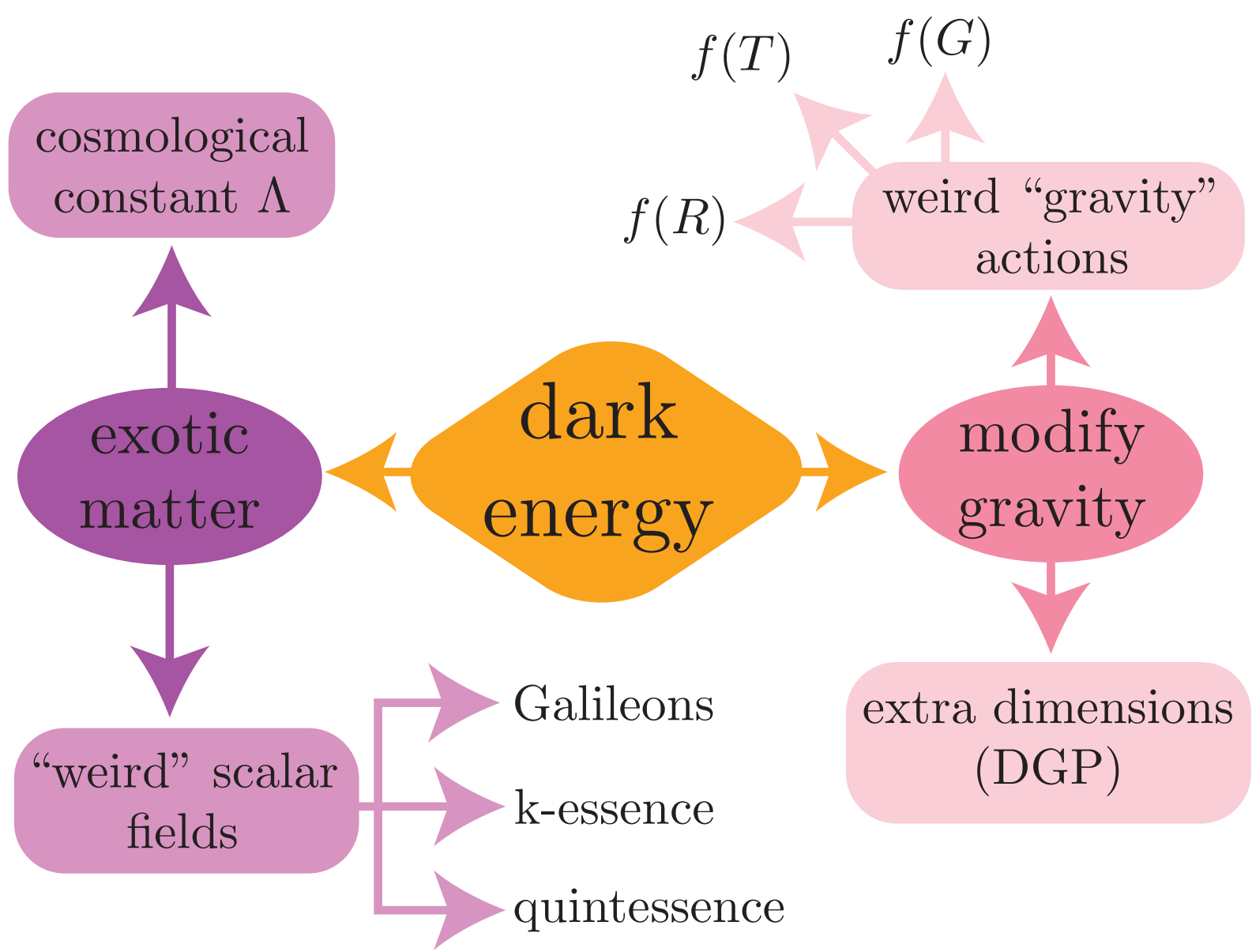
Other ideas

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Other ideas

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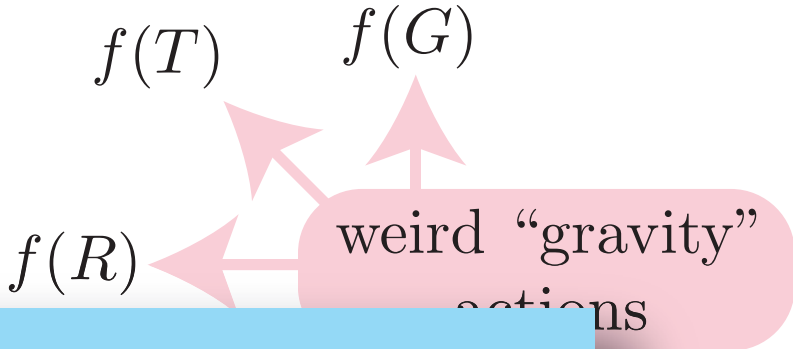
cosmological constant Λ

exotic matter

“weird” scalar fields

need more data on how dark energy evolves and clusters (goal of Euclid, BOSS, Dark Energy Survey...)

- Galileons
- k-essence
- quintessence



extra dimensions (DGP)

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Where do we stand?

Assessment

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- want to learn more? apply to grad school at UNB ...